Occultation Newsletter

Volume V, Number 10 January, 1993 ISBN 0737-6766

Published quarterly by the International Occultation
Timing Association
Joan B. Dunham-Editor

Occultation & Newsletter

Volume V, Number 10

January, 1993

ISBN 0737-6766

Occultation Newsletter is published by the International Occultation Timing Association. Editor: Joan Bixby Dunham; 7006 Megan Lane; Greenbelt, MD 20770-3012; U.S.A. Please send editorial matters to the above. Send new and renewal memberships and subscriptions, back issue requests, address changes, graze prediction requests, reimbursement requests, special requests, and other IOTA business, but not observation reports, to: Craig and Terri McManus; 1177 Collins; Topeka, KS 66604-1524; U.S.A.

FROM THE PUBLISHER

For subscription purposes, this is the fourth issue of 1992. It is the tenth issue of Volume 5, IOTA annual membership dues, including ON and supplements for U.S.A., Canada, and Mexico \$25.00 for all others 30.00

Annual IOTA membership dues may be paid by check drawn on an American bank, money order, cash, or by charge to Visa or MasterCard. If you use Visa or MasterCard, include your account number, the expiration date, and your signature.

ON subscription (1 year = 4 issues)

for U.S.A., Canada, and Mexico	20.00
for all others	25.00

Single issues are 1/4 of the price shown.

Although they are available to IOTA members without charge, nonmembers must pay for these items:

Local circumstance (asteroidal appulse) predictions	1.00
Graze limit and profile predictions (per graze)	1.50
Papers explaining the use of the predictions	2.50

Asteroidal occultation supplements will be available at extra cost: for South America via Ignacio Ferrin (Apartado 700; Merida 5101-A; Venezuela), for Europe via Roland Boninsegna (Rue de Mariembourg, 33; B-6381 DOURBES; Belgium) or IOTA/ES (see below), for southern Africa via M. D. Overbeek (Box 212; Edenvale 1610; Republic of South Africa), for Australia and New Zealand via Graham Blow (P.O. Box 2241; Wellington, New Zealand), and for Japan via Toshio Hirose (1-13 Shimomaruko 1-choine; Ota-ku, Tokyo 146, Japan). Supplements for all other areas will be available from Jim Stamm (117891 N. Joi Drive; Tueson, AZ 85737; U S.A.) for \$2.50.

Observers from Europe and the British isles should join IOTA/ES, sending DM 40.— to the account IOTA/ES; Bartold-Knaust Strasse 8; 3000 Hannover 91; Postgiro Hannover 555 829 - 303; bank-code-number (Bankleitzahl) 250 100 30.

IOTA NEWS

David W. Dunham

Membership Roster: It is time to update IOTA's membership roster; a new edition is planned for distribution with issue 12 in late April. By the end of March, you should communicate any updates to the McManuses in Topeka to their address in the masthead. We especially want to include telephone and fax numbers, and e-mail addresses, of IOTA members and ON subscribers, to help with rapid dissemination of time-critical events such as astrometric updates of asteroidal occultation predictions. We have telephone numbers for most members from returned information forms; let us know if you do not want your phone number published in the roster. Updates can be sent by e-mail to the McManuses, as described in the next paragraph.

New Electronic Mail Addresses: On p. 224 of the last issue, I gave my Internet e-mail address. I have learned that a more robust address for outside users is: david_dunham@jhuapl.edu

This avoids specification of my host computer for e-mail, which apparently is not known to some systems (a few users never were able to reach me at the previous address, although I could send messages to them). With both my first and last names specified, the e-mail system at Applied Physics Laboratory can find my host computer to route the message. If for some reason that doesn't work, it prints the message and it is sent to me by internal mail. The McManuses can also now be reached by e-mail; from Internet, their address is: 570-0611@mcimail.com

Either I.O.T.A. or CMcManus can be used in place of

the number preceding the "@", but the number is a more robust specification.

FAX We have purchased a FAX/Modem for our home computer. We expect to use it primarily for sending FAXes, but we can, by pre-arrangement, receive them. The FAX/modem will be using the same telephone line as the occultation line answering machine (301,474 4945). We do not have a way of sharing the use of the answering machine, and FAX/modem on that line except manually. We were offered some hardware that will monitor the type of call coming in and turn on the computer if it is a modem or a FAX, but that doesn't seem particularly practical for our machine. Our machine, a ZEOS 486/66, takes nearly a minute to start up when turned on, by which time most FAXes trying to send a FAX would have given up. We are still testing our FAX/modem. The modem part works fine. We wanted to use the FAX with WinFax Pro 3.0, but the installation program supplied with that software crashes before it finishes.

North American Asteroidal Occultation Supplement for 1993: Predictions of asteroidal occultations visible from North America during 1993, generated by Edwin Goffin, with events selected by Jim Stamm and charts annotated by David Werner, were distributed separately early this month, before the first event, rather than with the last issue, as stated in that issue.

Graze Supplements for 1993: Eberhard Riedel in Germany has used the coastline/border data supplied by me, and his own graze calculations, to generate maps and tables for 1993 hemispheric grazing occultation supplements in a form very similar to those that I have produced during the past few years. The tables do not include Z.C. numbers or double-star data, so I plan to write small programs to read the tables (that were supplied as ASCII files on disk) to add these data as soon as this issue of ON is completed. You should be receiving the graze supplement for your hemisphere within about two weeks, probably with the next issue, as noted at the end of this article.

Use of the Hubble Space Telescope (HST) by Amateur Astronomers: A news note inviting amateur astronomers to submit proposals to use HST was published on p. 14 of the February issue of Sky and Telescope. The annual April 30th deadline was mutually agreed upon by Steve Edberg, chairman of the (HST) Amateur Astronomers Working Group, and the Space Telescope Science Institute. A form and

useful information for writing amateur astronomer HST proposals are available from the Amateur Astronomers Working Group; c/o AAVSO; 25 Birch St.; Cambridge, MA 02138.

Change in ON Production: IOTA member Tony Murray in Georgetown, GA, said that he could duplicate ON at the print shop where he works, and include a cover, for less than we can duplicate it in this area, so this issue will be copied by him. Joan and I thank him for these efforts. Also, this issue will really be distributed from Topeka, KS, rather than from Greenbelt, MD; we thank the McManuses for their work with this.

Next Issue: The main purpose of this issue is to provide IOTA's information about planetary and asteroidal occultations during 1993. Extra work of updating predictions of occultations by (4179) Toutatis. other work, and the holidays prevented me from preparing the 1993 asteroidal occultation data before the start of 1993. The next issue will be distributed in about two weeks and will include important but not as time-critical articles and information that were not included in this issue. Unfortunately, just before printing this issue, the hard disk controller on our new PC misbehaved. Joan managed to recover the critical ON files and print this issue, but we will not be able to optimize the page layout of this issue as well as we would like. Because of the volume of material and our delayed schedule, we are effectively breaking up this issue into two issues, so that the next issue will put us approximately back on schedule. The issue after the next one, that is, #12, will contain information about occultations during the June 4th total lunar eclipse. If you plan to submit an article for that issue, we should have it by the second week of April.

CORRECTIONS TO ON 5, (5) and (9)

On p. 243 of the last issue, the 3rd sentence of the first full paragraph in the right-hand column should be "SAO 76225 should read SAO 76255". Also, 2 lines above that entry in the table on p. 118 of ON 5 (5), SAO 975990 should read SAO 075990.

A SOURCE FOR TIMEKUBES

Bob Nederman

As mentioned on p. 223 of the last issue, Tandy Corporation is no longer distributing the easy-to-use time source, the Weatheradio-Timekube, and by now has probably sold off all their remaining units in stock. My company, Astronomical Innovations; P.O. Box 14853; Lenexa, KS 66285, has a supply of 500 of these astronomically useful radios. We are selling them for \$22.95 plus \$2.05 for shipping, for a normal order of \$25 per unit, substantially less than Radio Shack's normal (\$39.95 pre-close-out) retail price; a check or money order should be in USA dollars payable to Astronomical Innovations. Orders can be placed by telephone at 1-913-894-5775. Inquire about overseas shipping, which will generally add about \$5 to the cost. Our supply should satisfy the astronomical community for 2 or 3 years, after which time Tandy Corporation may start selling the Weatheradio--Timekube, or an equivalent item, again. Until recently, Tandy Corp. did not realize that there was a continuing need for Timekubes in the astronomical community.

SOLAR SYSTEM OCCULTATIONS DURING 1993

David W. Dunham

General: My predictions of occultations of stars by major and minor planets, and by two comets, for 1993 are given in two tables whose contents are described in my articles about predictions of Solar System occultations for 1991 and 1992 in ON 5 (2) and in ON 5 (6). Most of the asteroidal occultation prediction material distributed by IOTA was prepared by Edwin Goffin in Belgium and is discussed in the third section. Sources of the predictions, other information, including stellar diameters (when significant), and notes about individual events, are given in the last sections.

For 1993, my annual Sky and Telescope article on planetary occultations was published in the February issue, pp. 76-77. Since that article is now limited to only North American events, and no good occultations of major planets occur there, the article was entitled

"Asteroid Occultations for 1993". Since there were no events in January, and Sky and Telescope's January issue was full, they decided to run the article in their February issue. ON was referenced for events outside of North America.

Reporting Observations: Reports of observations of any of these events should be sent to Jim Stamm; 11781 N. Joi Drive; Tucson, AZ 85737; U.S.A. (see his article elsewhere in this issue). Report positive or negative observations made under good conditions, but clouded-out attempts need not be reported. definite occultation is seen that could use some analysis for comparison with others, also send copies of the report to me at 7006 Megan Lane; Greenbelt, MD 20770; U.S.A., and to the chairman of the International Astronomical Union's (I.A.U.) Commission 20 Working Group on Predictions of Occultations by Satellites and Minor Planets, who is Lawrence Wasserman; Lowell Observatory; Mars Hill Road, 1400 West; Flagstaff, AZ 86001; U.S.A. Alternatively, observers may send their reports to their local or regional coordinators, who can then send the results to Stamm, and, when appropriate, to Lowell Observatory. The addresses of the regional coordinators are given in "From the Publisher" on p. 255 of this issue. Forms for reporting the observations can be obtained from Stamm or from the regional coordinators. indicate on the forms to whom copies are being sent. These forms are preferred, but the forms of the International Lunar Occultation Centre (ILOC), or the equivalent IOTA/ILOC graze report forms, can be used for reporting timed occultations or appulses. The main difference from reporting lunar events is that the name of the occulting body should be written prominently at the top of the form, and the report should be sent to neither ILOC in Japan nor to Richard Wilds. Also, if the asteroid is visible, the time that it merged with the star to form one apparent object, and the time the two were again noticeably separated, should be reported, with an estimate of whether the asteroid passed north or south of the star, if possible. Copies of the ILOC forms can be obtained from ILOC, from the IOTA secretary-treasurer (the McManuses in Topeka, KS), or from Richard Wilds; 3630 SW Belle Ave.; Topeka, KS 66614; U.S.A.

Event Selection: I made computer comparisons of my combined catalog with ephemerides of all of the major planets, comet P/Swift-Tuttle, the giant comet

P/Schwassmann-Wachmann 1 (P/Sm-Wm-1), and all minor planets for which Edwin Goffin predicted (see section below) at least one event under the selection conditions that we used for the main part of the North American Asteroidal Occultation Supplement for 1992: The star must be brighter than mag. 12.6; the magnitude drop must be at least 0.5; and for angular diameters smaller than 0".021, the star must be brighter than mag. 5.1; 0".021 to 0".050, brighter than mag. 6.1; 0"051 to 0".060, brighter than mag. 7.1; 0".061 to 0".070, brighter than mag. 8.1; and 0".071 to 0".079, brighter than mag. 9.1. In a few cases, these conditions were violated, such as for interesting objects (mainly, unusual light curves that may indicate duplicity) like 44 Nysa, 288 Glauke, 624 Hektor, 1220 Crocus, 2060 Chiron, 3123 Dunham, 5145 Pholus, and the two comets mentioned above. In a few cases, stars just slightly fainter than these limits were accepted when Goffin's prediction indicated that the path might pass over areas with large numbers of observers. The numbers of the minor planets included in my combined catalog searches included 2-4, 8-13, 15, 16, 18-20, 24, 27, 30, 31, 44-46, 49, 51, 52, 56, 58, 59, 70, 75, 78, 80, 85, 87-9, 97, 105, 107, 114, 141, 144, 146, 156, 171, 176, 181, 183, 203, 206, 216, 227, 236, 238, 258, 288, 303, 304, 324, 354, 357, 358, 407, 409, 410, 426, 444, 449, 451, 498, 511, 521, 532, 554, 566, 596, 624, 638, 654, 680, 704, 709, 712, 735, 772, 776, 895, 910, 1220, 2060, 3123, and 5145. Most of these asteroids were selected because occultations by them had been found earlier by Goffin or by Lawrence Wasserman at Lowell Observatory. many of the asteroids numbered in the high hundreds, those mainly with angular diameters less than 0.08, the searches were not performed for the whole year but only for a period of a few weeks centered on the date of events found by Goffin and Wasserman. In addition, Fresneau Astrographic Catalog (FAC; contains stars to 13th magnitude from declinations +4° to +32°) comparisons were made for 3, 10, 45, 52, 87, 107, 146, 511, 2060, 3123, 5145, and P/Sm-Wm-1. No FAC searches were done for some interesting objects simply because the ephemeris of the object remained outside the declination range of the FAC during all of 1993.

Note that I Ceres was not included in the searches. There is one 1993 prediction of an occultation by Ceres listed in my article in ON 5 (8), p. 205, where

a note discusses the extreme difficulty of the event, such that there is no need to include it in the main 1993 list in this issue.

Asteroidal Occultation Predictions by E. Goffin: The 1993 Asteroidal Occultation Supplement for North American Observers, prepared by Edwin Goffin with finder charts annotated by David Werner, were distributed separately early this month for IOTA members and ON subscribers in North America. Copies of Goffin's predictions and charts applicable to other parts of the world were sent by Jim Stamm a few months ago to regional coordinators for distribution to members and subscribers in their regions. For his 1993 predictions, Goffin converted his software to J2000 and used the new Positions and Proper Motions (PPM) J2000 star catalog, augmented with some other catalogs such as the FK5 and my version of the combined Lick-Voyager catalogs converted by him to J2000, rather than my Combined Catalog (CC). For a few asteroids, Goffin also used my version of Fresneau's Astrographic Catalog (FAC) that he converted to J2000. In a few cases, Goffin found occultations of PPM stars that are not in CC. For these, I converted Goffin's J2000 positions to B1950 and manually edited datasets to compute these events, usually successfully. In spite of the different catalogs and systems (my software and catalogs are still B1950), most of our predicted events are in common, and our predicted paths for the common events are generally (but not always) in good agreement. Consequently, we need to publish only a few finder charts in the regular issues of ON, since they have already been distributed with Goffin's predictions. In a few cases, we may publish 1° charts for some of the more crowded star fields on Goffin's charts, to facilitate locating the star to be occulted (the "target star"). These will be published alone, to be used in conjunction with Goffin's broader--field charts. Remember that the 1° charts are generated mostly from FAC. They are not needed as much for 1993 as they were for 1992, since Goffin has blown up many of the crowded fields of his plots to prevent this problem.

Comparison with the True Visual Magnitude Atlas (TVMA) often shows that some FAC stars are brighter, fainter, or very faint relative to their plotted magnitudes, indicated with B, F, or VF, respectively. "N" indicates that the star is not shown in TVMA.

Of the events found by Goffin that I tried to com-

pute, I failed for only one occultation, an occultation of 6.4-mag. PPM 574869 (= SAO 207178) by 680 Genoveva that Goffin predicted for Australia on August 24th. My calculation shows that the closest approach will occur a day earlier with the asteroid missing the star by 4'. We both used orbital elements from MPC 16391, so this discrepancy remains unexplained.

Most of the PPM stars have SAO numbers, which I prefer to use, considering the more widespread availability of the SAO catalog. Also, Goffin assigned sequential numbers to some of the catalog sources. including the FAC, where the stars remain unnumbered in my version. For the Lick-Voyager catalogs, DM numbers are often given (especially for L 3 and L 5 (Lick-Uranus and Lick-Neptune) stars in Sagittarius and Capricornus. Goffin only gave the four least significant digits of the DM numbers of these stars, most of which are from the Cordoba Durchmusterung (all Lick DM numbers south of -22°), where the numbers are all in the 10,000's for stars in Sagittarius and Capricornus. So 10,000 needs to be added to the DM numbers for these stars in Goffin's predictions, for example, for most of the stars occulted by 24 Themis.

Explanation of Data in Tables 1-3: A complete explanation of the data in Table 1, and a partial explanation (actually, covering most of it) of the data in Table 2, was given in my article, "Solar System Occultations during 1991", in ON 5 (#2, December 1990), starting on p. 39. The explanation of the rest of the Table 2 data was given in my article, "Solar System Occultations during 1992", in ON 5 (#6, December 1991), starting on p. 132, and an explanation of Table 3 starts on p. 133 of the same article.

Local Circumstance/Appulse Predictions: Joseph E. Carroll; 4261 Queen's Way; Minnetonka, MN 55345; USA, computes the IOTA appulse predictions for all IOTA members. Note that the star source code logic of this program has not been updated, so that the source codes in the appulse predictions will sometimes differ from that given under \underline{S} in Table 2 described above. In case of disagreements, use the Table 2 code. Hans-Joachim Bode distributes similar predictions to IOTA/ES members. The format of these predictions is nearly self-explanatory and contains virtually all of the information that an observer needs. Columns headed \underline{D} and \underline{S} following the SAO number

give the double star code and star position source code (but see the remark above), respectively. Next are the star's DM/ID No., then the star's MAG (visual mag.), OCC. DMAG (occultation magnitude drop), and DUR SEC (central occultation duration in seconds). This is followed by the U.T. and distances (in arc seconds, kilometers on the sky plane, and in terms of object diameter) of local closest approach. The distances are positive if the asteroid passes north of the star (this means that the path would be south of the observer's location). The elongation (ELG, angular distance from the star) of the Sun and Moon are given, as is also the Moon's percent sunlit (PSNL).

World Maps: World maps by Mitsuru Sôma are published here only if the event is not included in Goffin's predictions; or if the star is mag. 8.0 or brighter; or if the star is double, and I have drawn a line showing the 2nd component path; or if there is more than about 0".5 discrepancy with Goffin's prediction; or if there is a recent astrometric update. The charts show the Earth as seen from the asteroid at the time of the event; the hatched curve marks the sunrise or sunset terminator, with hatches on the night side.

Regional Maps: The three regional maps showing quarterly Solar System occultations between latitudes +65° and -50° starting on p. 264 are like the ones for the Toutatis occultations starting on p. 237 of the last issue, rather than like the old quarterly maps, such as the ones starting on p. 251 of the last issue. That is, except for some of the occultations by major planets. they will have time lines at 2-minute intervals, but will not have longitude or latitude tick-marks or labels, and the paths will be hand-labelled. "Time lines" that seem out of sequence, or slant the wrong way, are moonrise or moonset lines. The enclosing rectangles will have the same latitudes and longitudes, and the plots are still false projections with horizontal and vertical scales both linear to facilitate plotting or measuring of coordinates. The charts cover the 4 months from February through May, since this issue is too late for the January events, and ON 5 (12) needs to be distributed in time for the June 4th lunar eclipse.

Finder Charts: Previously, I produced 3° and 1° charts for many events not predicted by E. Goffin. 1 have not had time to regenerate this capability since losing access to the necessary hardware last September. I plan to produce unlabelled star charts, which can be manually labelled, in time for the next issue.

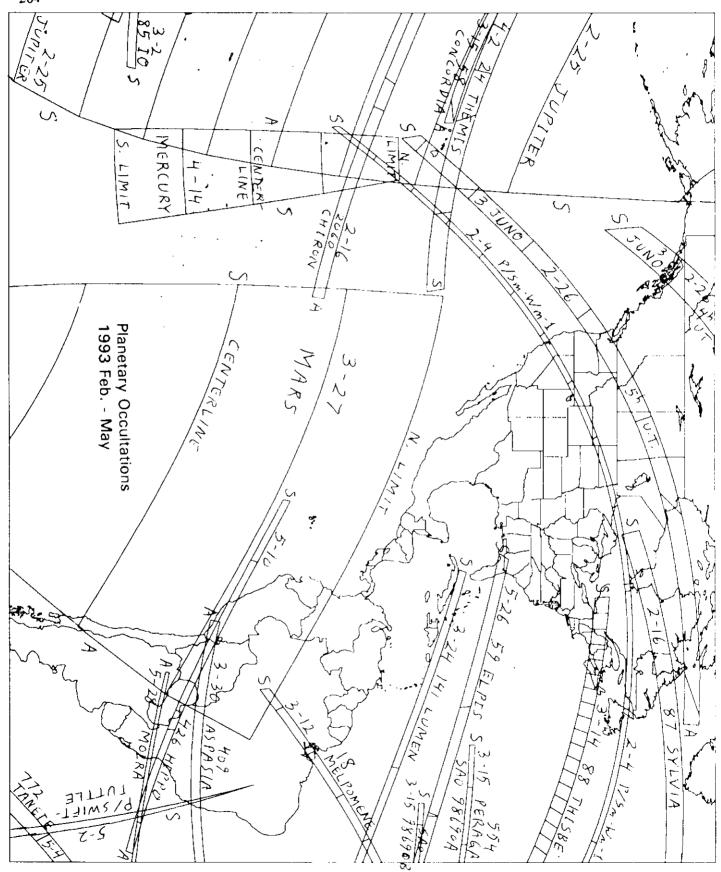
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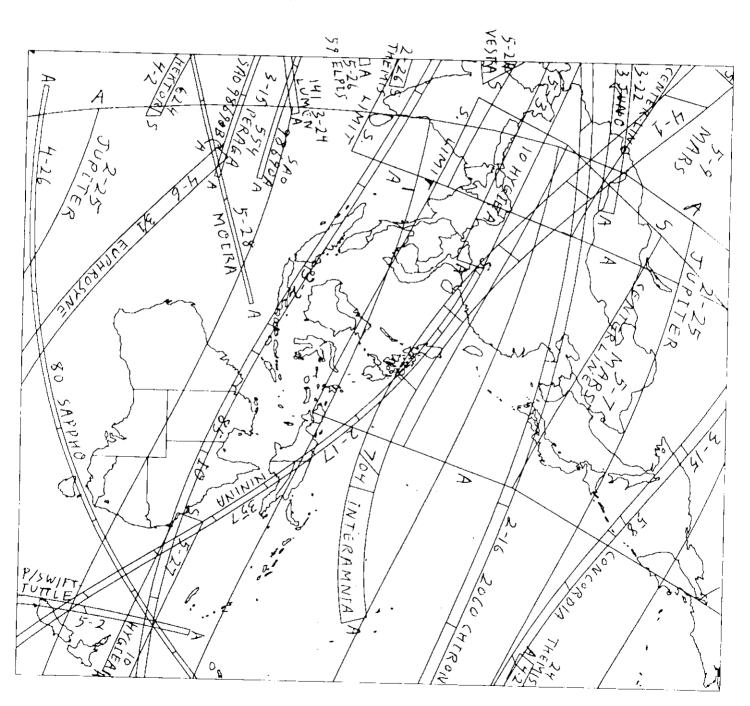
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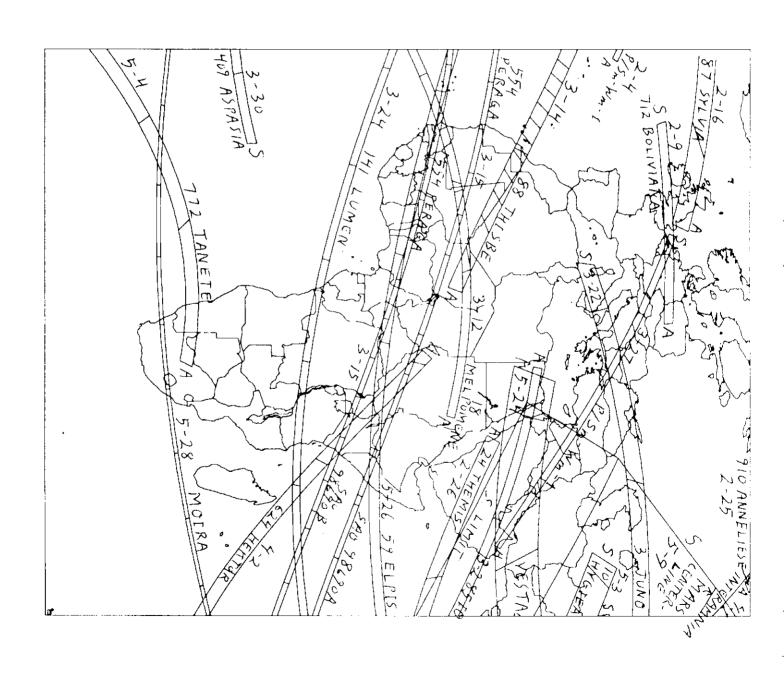
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Planetary Occultations 1993 Feb. - May



Planetary Occultations 1993 Feb. - May



<u>Priority List:</u> A priority list of events most likely to have last-minute astrometric updates will be given in the next issue.

Occultations by Major Planets: I have included all occultations of major planets given by L. Wasserman, E. Bowell, and R. Millis in Astronomical Journal (AJ) 103 (103), p. 2089, as well as a few additional events that my search revealed. The occultations by Jupiter will be very difficult; I included them only because they were listed by Wasserman et al. No occultations by Saturn were found, but some very difficult events, perhaps visible in infrared bands with large telescopes, are given by Bosh and McDonald in AJ 103 (103), p. 983. Similarly, some difficult occultations of faint stars by Uranus and Neptune are given by D. Mink and A. Klemola in AJ 102, p. 389. Possible occultations by Pluto or by Charon are listed by D. Mink, A. Klemola, and M. Buie in AJ 101, p. 2255. One of the best Pluto occultations of the rest of this decade, involving a 12.4-mag, star possibly visible from Japan or eastern Australia, is predicted to occur on October 3rd. Wolfgang Beisker and some other members of IOTA/ES are making plans to try to observe the event from Australia. Unfortunately, the small elongation from the Sun will make last-minute astrometry difficult, but Pluto's motion is so small that a good prediction may be possible from plates taken a few months before when the elongation is larger.

Notes about Individual Events: No notes are given for events in January, since this issue will unfortunately be distributed after those events.

Feb. 4: This is the giant periodic comet Schwassmann-Wachmann 1, in a nearly circular orbit beyond Jupiter; its diameter is only a guess. In December, the comet underwent an outburst, and in late January, it was still brighter than usual, but still much fainter than the star. Dimming in the coma may occur within one or two hundred km of the path, whose location is quite uncertain due to the AC source for the star's position and the object's relatively large distance from the Earth (so the event might occur anywhere in North America). A recent position for the star will probably be obtained before the event to improve the prediction. I will try to make a finder chart for the star before the event, but distribution of it to potential observers will be a problem, and may be possible only by FAX. Feb. 16, Chiron: Chiron is also apparently a giant comet nucleus, although at its greater distance, it is not as active as Schwassmann-Wachmann 1. Like the Feb. 4 event, the prediction is very uncertain due to the AC source for the star's position, so an update of the star would again help. The world map shows that the event could be visible almost anywhere in the Pacific Ocean, eastern Asia, or Australia.

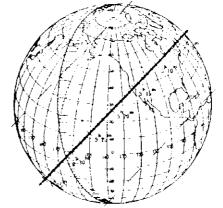
<u>Feb. 25: Jupiter</u> will have a negligible 0".12 defect of illumination.

Feb. 26, Juno: The two occultations by Juno on this date are predicted by E. Goffin to occur over 1" southeast of my paths. The first event, with a very small magnitude drop, would normally not be included, but it is presented here because of the other more favorable event an hour later.

Notes about individual events after February will be given in the next issue.

Table 3. Stellar Angular Diameter Information

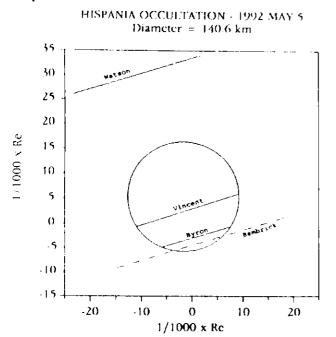
1993 Date	PLANET NO. NAME	STARD SAO/DM/Id	
Feb 17	357 Ninina	96403	1.44 2471 289 ^{MS} 8.1
Mar 13	88 Thisbe	96656	0.28 546 266 1.7
Mar 24	141 Lumen	156969	0.67 1080 70 3.7
Apr 2	624 Hektor	211153	0.67 1080 70 3.7 0.23 847 53 1.9
May 26	59 Elpis	140990	0.64 896 70 3.3
May 28	638 Moìra	159886	0.61 574 65 2.5
Juni 15	20 Massalia	184751	0.64 896 70 3.3 0.61 574 65 2.5 0.74 924 77 3.6 1.89 3524 90 11.1
Jun 21	89 Julia	75188	1.89 3524 90 11.1
Jul 17	776 Berbericia	943320	U.YI ZIDI DZ 0.1
Aug 9	89 Julia	56655	1.36 2149 81 7.4
Aug 12	303 Josephina	146585	5.94 9127 974 31.8
Aug 13	171 Ophelia	79782	1.97 5376 120 14.0
Aug 17	304 Olga	128584	0.53 371 51 1.9
Aug 22	27 Euterpe	+ 2° 131	0.27 293 152 1.2
Aug 31	19 Fortuna	L 5 1485	0.18 190 53 0.8
Oct 1	183 Istria	214187	1.82 1548 198 7.3
Oct 10	206 Hersilia	94649	0.22 327 61 1.1
Oct 15	30 Urania	78016	0.48 536 66 2.2
Oct 17	444 Gyptis	93255	0.78 817 87 3.4
Oct 19	532 Herculina		0.48 536 66 2.2 0.78 817 87 3.4 0.19 365 88 1.1
Nov 9	56 Melete	118362	0.85 2049 80 5.7
Nov 13	358 Apollonia	165042	0.75 1232 122 4.2
Nov 27	444 Gyptis	110612	0.55 620 90 2.5
Nov 30	144 Vibilia	+24° 1499	0.37 416 58 1.7 0.67 812 73 3.2
Dec 26	203 Pompeia	79240	0.67 812 73 3.2



Anonymous by P/SM-WM-1 93 Feb 4

OCCULTATION OF SAO 181281 BY (804) HISPANIA ON 1992 MAY 5

This occultation was observed by Keith Vincent at Havelock North, New Zealand (duration 11.4 seconds); Jeff Byron, Sydney, N.S.W., Australia (duration 7.5 seconds); and Colin Bembrick, Bathurst, N.S.W. (duration uncertain, 2.5 to 5.5 seconds); details are given in Graham Blow's article on pages 13 and 14 of Circular CQ 92/2 (September 1992) of the Occultation Section of the Royal Astronomical Society of New Zealand, from which the figure below is reproduced.



ASTEROID (MAINLY TOUTATIS) NEWS

David W. Dunham

Predictions of occultations by (4179) Toutatis during its close approach to the Earth were given in an article starting on p. 233 of the last issue. None of the Toutatis occultations is known to have been observed. The weather was bad in most of North America for the events that occurred there, and I have heard of only a few appulse observations. I got the first orbit updated with radar data from Don Yeomans at the Jet Propulsion Laboratory (JPL) on December 7th, but I did not have time to make the changes needed to produce an

accurate ephemeris for an object passing so close to the Earth. By about December 12th, we worked out an arrangement where Don sent me a B1950 ephemeris by e-mail, and I was able to generate updated predictions with these data. The last updated ephemeris came on the 18th, but already by the 12th, the radar-updated ephemeris was so good that the new radar observations showed that Toutatis was within about 1 km of its predicted distance. However, the cross-track error might be a little larger, so that successful occultation observations might be useful to supplement the radar data to provide accurate 3-dimensional data for Toutatis that could be used to refine the predictions for future close approaches by this object (an improved position for the star, which is expected from Hipparcos satellite observations now in progress, would be needed for this). The updated predictions were near the centers of the uncertainty zones that I plotted on the three maps in the last issue. The only practical way to distribute the updated predictions before the holidays was by e-mail, and I sent predictions to about 30 observers around the world that way. John Spencer at Lowell Observatory copied some of my data to a larger e-mail list of mainly professional astronomers coordinating astronomical observations of Toutatis, and this established some new contacts that will be useful for future asteroidal occultations.

By late December, current star position errors seriously magnified the width of the uncertainty zones as Toutatis rushed away from the Earth. These errors (half-widths) are given in km in the table below:

Cat. soc	irce 0	or S	A R	orU	Ł	5	Hipparcos
Error							0.002
1992	Dist.						
Dec. 8	0.024	18	11	5	2	1	0.04
Dec. 15	0.050	37	22	11	4	2	0.07
Dec. 23	0.100	74	44	22	8	4	0.15
Dec. 31	0.155	114	68	34	12	6	0.23

The star catalog source codes, or "Cat. source" in the table, for occulted stars were given under "S" in the 2nd table on p. 235 of the last issue. C = Astrographic Catalog, S = SAO, A = AGK3, R = AGK3R, U = USNO Zodiacal Zone (Z.Z.), L = PPM high-precision subset, and 5 = FK5. The distance of Toutatis in astronomical units is listed after the 1992 date. Since the radar data (spectacular radar images have been published in Science News, Space News, and elsewhere) show that the long axis of Toutatis is about 5 km, the problem of spacing enough observers

across the uncertainty zone to really catch the occultation is apparent with current star positions. Hopefully, when Toutatis makes an almost identical flyby in December 1996, Hipparcos (specifically, Tycho catalog) positions accurate to a few milliarcseconds will be available for most of the stars so that observers can be precisely positioned to catch the occultations. I presented the above material at a workshop on first results of the Toutatis flyby that was held at the University of Arizona in Tucson on January 4th, just before a meeting on hazards due to comets and asteroids that I also attended.

JPL's Steve Ostro recently showed me the impact that upgrades to the radars at Goldstone and Arecibo during the next two years can have on asteroid science. These will allow crude imaging of virtually all mainbelt asteroids larger than 200 km, and many of the inner-belt objects over 100 km. If enough observations can be scheduled, the upgraded radar data will also allow precision orbits to be determined for these asteroids. The better orbits coupled with good star positions determined from Hipparcos observations should make possible accurate prediction of occultation paths months and years in advance, without the need for last-minute astrometry.

Steve Ostro has said that the radar observations of Toutatis, which have a distance binning of 19 meters, have revealed no satellites or debris clouds. might be expected for an asteroid that has probably made previous very close approaches to the Earth that would tend to perturb any possible satellites away from the week gravity of Toutatis. Nevertheless, on January 18th, Petr Pravec in the Czech Republic made a remarkable discovery - he found an object about 2.5 magnitudes fainter than Toutatis about 40" away and moving with a similar motion. The object was about 90" away the next night, in spite of Toutatis' daily motion of over 1000". However, Alan Stern soon pointed out that the gravitational sphere of influence of Toutatis would be at most 3". Brian Marsden solved the mystery when he found that another asteroid, 1992 YG3, that had been discovered in Japan on December 30th, did indeed pass near Toutatis in the sky on January 18th, and had a remarkably similar motion at that time.

CORRECTIONS TO GRAZING OCCULTATION PREDICTIONS

David W. Dunham

Northern-limits: During the last two years, darklimb northern-limit waxing-phase grazes, which generally occur from February through June, have usually been shifting south of IOTA's prediction version 80M and 80N by a few tenths of an arc second, and this is expected to continue during 1993. You should adjust your coverage to expect anywhere from no shift to a 0.75 southward shift of the shadow, using the vertical scale on the left side of your ACLPPP profile. I would recommend aiming for a location 0.3 south of the most interesting part of the profile (that is, the horizonal line that is closest, say within 0".2 of, the most predicted profile points), but, if this is within 0".7 of the highest mountain top on the profile, you should be sure that someone observers from 0".7 under the highest mountain, to give a high probability of seeing something during the graze, and not just a miss (no occultation). In general, with 3 or more stations, it is best to try to cover a full arc second of vertical distance, at least covering from highest mountain top to either 0".7 south of it, or 0".3 south of the most interesting region, whichever is farthest south.

Southern-limits: Dark-limb southern-limit waxing-phase grazes near the lunar South Pole, in particular, from Watts angles 165° to 180° when the profile points are plotted as numbers from 3 to 7, have also been shifting south, usually by 0".4 to 0".5. These events usually occur in the evenings from October through March. Fortunately, there are large mountains in this region, so that the danger of seeing a miss is much less than for the smooth features usually found in the northern polar regions. There is some evidence that most of the southern Cassini region, where the profile points are mostly plotted as numbers from 2 to 4, is all a little farther south (higher) than our current ACLPPP predictions indicate, so this correction may affect waning-phase grazes, also, to Watts angle 187°.

REPORTS OF ASTEROIDAL APPULSES AND OCCULTATIONS

Jim Stamm

If you do not have a regional coordinator who forwards your reports, they should be sent to me at: 11781 N. Joi Dr. Tucson, AZ 85737 USA. Names and addresses of regional coordinators are given in "From the Publisher" on Occultation Newsletter's front page. All times in this report are UTC.

I have summarized all of the reports that I have received for the last half of 1990 in the following two tables and section of notes. Table 1 lists the 1990 date, minor planet, occulted star, IDs of successful observers, and references to any notes. Table 2 lists the observer's ID, name, nearest town to location of observation, country (includes state or province for North America and Australia), and the total number of observations made in the period. The notes section details those events that included positive observations, or other significant information that could not be reported in the tables. I am not including notes on those observations that may have been spurious unless there is some sort of confirmation, or the fact that something may have happened is relevant to another observation. Instead, I will place an asterisk (*) in the Notes column to indicate that I have received a report with more than a "no event...." in it.

Table 1. Asteroidal appulses and occultations: Jul-Dec 1990.

199	0	(inor	Planet	Cat	Star	Observers M	otes
Jul	02	8	Flora	SAO	186885	SmcHut.AndD1k	
Jul	04	176	Iduna	ACK3	+16° 2602	BqsBulCanDssGb	£
						GrcIelMddNelRu	zVq1
Jul	06	6	Hebe	SAO	120195	SmcHut.And	•
Jul	09	39	Laetitia	SAO	119674	GrhGrt	
Jul	09	8	Flora	LickV	5980	Dik	
Ju!	10	196	Philomela	LickV	2003	SmcLap	
Jul	10	86	Semele	LickV	5597	CooOveMitVnbLag	Smc
Jul	12	598	Octavia	LickV	872	StaSmo	
Jul	14	224	0ceana	SAO	183675	Sta	
Jul	15	3	Juno	AGK3	-01° 1862	OveCooMit	1
Jul	19	679	Pax	LickV	6945	HozHon	
Jul	20	211	Isolda	SAO	109396	Dik	
Jul	28	8	Flora	ĎM	-22° 4543	AndSmcDik	
Ju]	28	8	Flora	LickV	2309	ProVnbAndSmc	
Jul	28	8	Flora	SAO	186216	DaeProVnb	
Jul	29	8	Flora	SAO	186209	Hon	
Jul	31	732	Tjilaki	AGK3	+01° 2810	BlkStgAndSmc	
Jul	31	454	Mathesis	SAO	209137	BlkLapCai	
Aug	09	10	Hygiea	AGK3	+00° 2899	DikStg	
Aug	09	679	Pax	SAO	186343		2
Aug	10	96	Semele	LickV	1633	BlkBrySmc	
Aug		516	Amherstia	SAO	189599	CooOveSmiKni	
Aug		38	Leda	ACK3	+00° 2856	CooKniOveSmiVni	5
Aug		145	Adeona	SAO	129413	Hut.Smc	
Aug		441	Bathilde	SAO	159572	Sta	
Sep		81	Terpsichore	AGK3	+29° 0648	WabKe1SamHonWe:	i
Sep		679	Pax	SAO	186284	GarTavLyzChl	
Sep		704	Interamnla	FAC	895087	And	
Sep	C6	377	Campania	AGK3	+09° 0045	Smi	
Sep		501	Urhixidur	SAO	190967	DaeProBlkLoaSmo	2 3
Sep		40	Harmonia	SAO	190308	Wid	*
Sep		276	Adelheid	AGK3	+12° 0511	Hak	
Sep		260	Huberta	LickV	5131	Sta	
Sep	11	20	Massalia	SAO	164484	Sta	

<u>Tab. 1</u>	(Cont.) Asteroida	l appul	ses/occulta	tions: Jul-Dec '90
Sep 16	121	Hermione	ĸ	2350	DwdHavSatSms
Sep 22	689	21ta	SAO	146303	CooOveWesMitSmi
Sep 24	19	Fortuna	AGK3	+22° 0623	DflDssKocksz 4
-					MeuMlnMosSzaZal
Sep 29	19	Fortuna	AGK3	+22° 0643	BilBnnBulDflDss
					ErnFauGbfHokHolKhl
					MeuMlnMosShtSzcVlr
Sep 29	160	Una	LickV	5951	Stg
Sep 30	51	Nemausa	SAO	163983	TamRbbDeaMonGea
					PryLyzChiWimJoh
Sep 30	451	Patientia	FAC	104734	Dik 5
Oct 02		Flora	AC.	11836	Sta
Oct 11	196	Philomela	Li ckV	6717	OveSmiBlnVnb
Oct 14	537	Pauly	SAO	189034	Ven
Oct 14	306	Unitas	SAO	128623	StaSmcAnd
Oct 14	494	Virtus	DM	+19° 0547	Sta
Oct 22	120	Lachesis	AGK3	+15° 0115	StaKrtSpr
Oct 22	185	Eunike	AGK3	+02° 1213	BlkSmcAnd 6
Oct. 22	139	Juewa	AGK3	+07° 0061	7
Oct. 24	127	Johanna	SAO	189449	BgsDssGcvGrcLooVgl
Oct. 30	804	Hispania	AGK3	+25° 1097	Hon
Oct 30	661	Cloelia	ACK3	+35° 0563	HonStaManPal
Oct 30	506	Marion	LickV	2928	CmbDflDssGen
Nov 12 Nov 15	60 704	Echo	AGK3	+13° 0263	PryStaManPal
Nov 15		Interamnia	FAC	885701	VenHonStaManPal
MOA 13	924	Toni	AGK3	+09° 0488	BrrDflFauFenFrb
Nov 17	216	Kleopatra	FAC	11 2	GloGrcPrcThm Rel B
Nov 19	537	Pauly	SAO	11.2 mag 189987	Rei 8 SmiJoo
Nov 20	838	Seraphina	AGK3	+10° 0029	SmiVnbAouDf
1407 20	030	Serapitina	nun-	+10 0029	lDssIelKkn
Nov 23	323	Brucia	SAO	193254	DikBikAnd
Nov 26	614	Pia	AGK3	+08° 0147	CooBin
Dec 04	31	Euphrosyne	AGK3	+05° 1819	Sta
Dec 05	701	Oriola	AGK3	+18° 0239	LapB1kSmc
Dec 05	107	Camilla	LickV	2528	DssGbfLan
Dec 09	704	Interamnia	FAC	900107	GrhHonStaSamWar 9
Dec 10	451	Patientia	FAC	87231	CanGro
Dec 14	17	Thetis	LickV	23215	Dss
Dec 15	121	Hermione		Anon.	And
Dec 15	121	Hermione		Anon.	And
Dec 19	451	Patientia	ACK3	+08° 0360	MamCvgDssGdiAnd *
Dec 20	860	Ursina	AGK3	+33° 0237	GeaStá
Dec 25	121	Hermione		Anon,	LapBlk
Dec 28	121	Hermione	FAC	188065	BulDflMos
Dec 31	205	Martha	DM	+04° 0190	AouDflDssHllMos

Table 2. Observers and locations of events: Jul-Dec 1990.

ID	Observer	Town	Country N	Ιo,
And	Anderson, Peter	The Gap	Queensland - AUS	12
Aου	Arnaout, W.	Kalaa Sghira	Tunisia	7
Bni	Baroni, Sandro	Milano	Italy	1
Brz	Barruezo, Jose.	Granada	Spain	1
Brh	Barthes, Jacques	Castres	France	1
Bff	Baruffetti, Pietro	Massa	ltaly	1
Bat	Bath, KL.	Emmerdingen	Germany	1
Bel	Bellot, Luis	Granada	Spain	1
Bor	Benier, Jacky	Varades	France	1
Bin	Bentlin,	Freiburg	Germany	1
Bln	Blane, D.	Henly on Klip	South Africa	2
Blk	Blanksby, Jim	Wandin	Victoria - AUS	8
Bsc	Blasco, Julian	2aragoza	Spain	1
Bnn	Bonninsegna, R.	Dourbes	Belgium	3
Brr	Borras, Vincente	Benicarlo	Spain	2
Ass	Bossalaers, Mark	Berchem	Belgium	1
Bgs	Bourgeois, Jean	Ciney	Belgium	4
Bi1	Bril, Henk	Urmond	Netherlands	2
Bry	Bryant, Ken	Langwarrin	Victoria - AUS	1
Bul	Bulder, Henk	2oetermeer	Netherlands	4
C11	Callens, B.	Gent	Belgium	1
Cai	Camilleri, Paul	Cobram	Victoria - AUS	1
Cns	Canales, Oscar	Pinsoro	Spain	1
Can	Candela, Bernard	Sollies-Pont	France	'3

VG	Cavagna, Marco	Sormano	Italy	2	Pnn	Pannier, Lutz	Gorlitz	Germany
hİ	Child, Jack	Table Mountain	California - USA		Prr	Pereira, Alfredo	Lisboa	Portugal
din	Colomba, Armando	Reggio Calabria	Italy	1	Pak	Pigulski, Andrzej	Wroclaw	Poland
00	Cooper, Tim	E. Rand/Malelane		6	Por	Porcel, Aniceto	Granada	Spain
1	Daiffallah, K.	Alger	Algeria	7	Pro	Porcini, Roberto	Salerno	Italy
e	Dale, S.	Pietermaritzburg		2	Ppl	Pouplier, Alphonse	Ciney	-
3	Dean, Fred	Victoria	Brit, Col CAN	1	Pro		•	Belgium
k	Dickie, Ross	Gore		7		Prosser, G.	Pietermaritzburg	
<u>.</u>			New Zealand		Pry	Pryal, Jim	Seattle/Easton	Washington - USA
	Dunham, David W.	Greenbelt	Maryland - USA	1	Blt	Ramon, Luis	Granada	Spain
s	Dusser, Raymond	Kalaa Sghira		1.2	Rsp	Raspadori, G.	Bologna	Italy
C	Ernst, Christoph	Graz	Austria	2	Rdn	Raudino, Paolo	Civitavecchia	Italy
v	Estever, Carles	Els Hostalets	Spain	l	Ribib	Robb, Russ	Victoria	Brit. Col CA
ı.	Faure, G.	Varces	France	3	Frb	Rodriguez, F.	Sevilla	Spain
r	Pederspiel, Martin	Freiburg	Germany	1	Ruz	Ruiz Fernandez, J.	Santander	Spain
d	Feldmann, JB.	Di jon	France	1	Ru1	Rulz Mazo, S.	Granada	Spain
1)	Fernandes, J.	Pedrogao Pequeno	Portugal	1	Sat	Salthouse, Andrew	Millington	New Jersey - US
1	Foglia, Sergio	Lacona	Italy	1	Sam	Samolyk, G.	Mtlwaukee	•
0	Gallo, Vicenzo	Salerno	Italy	1			Greenfield	Wisconsin - USA
С	Garcia, Joaquim	Lisboa	Portugal	5	Sau	Sauter, Christof	St. Margarethen	Switzerland
,	Garcia, Ruben	Monteviden	Uraguay	1	Sch		-	
n	Genovese, Marco	Torino				Schnabel, Carles	Barcelona	Spain
a			Italy	1	Shk	Schoenmaker, A.A.	Roden	Netherlands
	George, Anthony	Salem	Oregon - USA	2	Sht	Scholten, Alex	Eerbeek	Netherlands
P	Gigli, Paolo	Pistoia	Italy	2	Sch	Schwaenen, Jean	Marcinelle	Belgium
f	Gobet, Franck	Villefranche/\$.	France	3	Sm1	Smit, J.	Pretoria	South Africa
7	Goncalves, Rui	Lisboa	Portugal	2	Smc	Smith, Charlie	Woodridge	Queensland - AU
'n	Gracias, Nuno		Portugal	1	Sms	Smith, Scott	Fultondale	Alabama - USA
h	Graham, Frances	East Pittsburg	Penn USA	2	Sp1	Spell, Jerzy	Walbrzych	Poland
	Graham, Theresa	East Pittsburg	Penn USA	1	Spr	Springob, C	Slegen	Germany
i	Gualdoni, Carlo	Sormano	Italy	1	Stg	St. George, Lou	Auckland	New Zealand
h	Guenther, Elke	Freiburg	Germany	1	Sta	Stamm, Jim		
Ż	Harvey, Roger	Concord	N. Carolina - USA			· ·	Tucson	Arizona – USA
	Hauk, Robert	Portland			Sut	Sutterlin, Peter	Freiburg	Germany
			Oregon - USA	1	Sza	Szabo, Sandor	Boly	Hungary
ı	Holler, Gert	Graz	Austria	1	Szc	Szolcsanyi, Gyorgy	Pilisszen	Hungary
K.	Holler, Klaus	Graz	Austria	2	Szl	Szollosi, Attila	Kecskemet	Hungary
1	Hollis, Andrew	Northwich	United Kingdom	1	Tam	Tatum, Jeremy	Victoria	Brit. Col CA
7	Holtz, John	Russellton	Penn USA	1	Tav	Taverez, Gabriel	Montevideo	Uraquay
n	Honkus, Edward	Carnegle			Tsl	Tesi, Luciano	Pistoia	Italy
		Russellton	Penn USA	7	Thun	Thomas, S.	Aix-en-Prov.	France
t	Horst, Schmidt	Freiburg	Germany	1	Tso	Tissot, M.	Villeneuve d'A.	France
t	Hut cheor	Sheldon/Warwick	•	3	Trl	Torrell, Sebastia	Barcelona	Spain
1	Tela, Antonio	Reggio Calabria	Italy	3	Vbq	Van Billegoy, E.	Drut en	Netherlands
r.	Johnson, Randy	Seat.tle	Washington - USA		Vnb	Van Blommestein,P		
o	Jooste, J.	Reitz	-				Simon's Town	South Africa
z	•		South Africa	1	Vq1	Van Gestel, Jan	Ceel	Belgium
	Kasz, Laszio	Boly	Hungary	1	Loo	Van Loo, E.	Ciney	Belgium
ì	Keith, Lee	Мееле	Wisconsin - USA	1	Vso	Van Soom, H.	Sent	Belgium
m	Klemencie, R.	Corenja Vas	Yugoslavia	1	Ven	Venable, Roger	Wrens	
j	Kright, J.	East Rand	South Africa	2			- Hard Labor Cr. S.	P.
(2	Kocsis, Antal	Balatonkenese	Hungary	1			Ft. Gordon	Georgia - USA
d	Kohl, Ferdinand	Ust er	Switzerland	1	Vid	Vidal, Joaquin	Monegrillo	Spain
1	Kohl, Mike	Wald	Switzerland	2	Vlr	Voller, Wolfgang	Graz	Austria
	Kosa-Kiss, Attila	Salonta	Romania	1	War	Waraczynski, Sally	Muskego Co. Park	
į.	Kosir, B.	Ljubljana	Yugoslavia	1	Wab		-	
	Kretlow, Mike					Warner, Brian	Colorado Springs	
		Stegen The Basin	Germany	2	Wei	Weier, Davld	Brooklyn	Wisconsin - USA
	Larkin, Patricia	The Basin	Victoria – AUS	5	Whk	Weith-Knudsen, N	Tisvildeleje	Denmark
	Laurent, Dirk	Gent	Beigium	1	wes	west, D.	Lanseria	South Africa
ì	Le Guern, Vincent	Villeneuve d'A.	France	.3	Wid	Widdop, H.J.	Plerrefonds	Quebec - CAN
ì	Loader, Brian	Black Birch			Wim	Williams, Ernie	Table Mountain	California - US
		Christchurch	New Zealand	1	₩Is	Wils, P.	St-CatWav.	Belgium
2.	Lyzenga, Greg	Table Mt.	California - USA	2	Whib	Wubbena, E.K.	Oosterhout	Not her lands
T.	Majumdar, T. K.	Calcutta	India	1	Zal	Zalezsak, Tamas	Balatonkenese	Hungary
7	Manly, Peter	Central	Arizona - USA	3				,
i	Marti, Josep	Mataro	Spain	1				
	Marx, Harald	Stuttgart	Germany	ì	NOTE:	5 (Jul-Dec 1990);		
	Mazalrey, Jean	Vernon	Erance	1		17701.		
z .	Mazalrey, Pierre	Vernon					_	
			France	1	l .	Jul (Jun) 15 .	Juno. See (O.	N. 5(4), p.9
h •	Michon, Jean-Pol	Herment	France	1		This is a corre-	cted date.	
3	Middleton, R.W.	Brightling, Col.		1	2			- 03:
i.	Mitchell, H.	Pennington	South Africa	3			See $\{0, N, 5(4)\}$	
n,	Montoya, Mike	Mariposa	California - USA	1		Observers were:	BelBatBffBgsBl	nBltBn!BnnBn
	Moretti, Stefano	Forli	Italy	1		BnrBrhBrrBrzBsc		
t	Morillon, Eric	Liguge	France	2			-	
:			Algeria	4		FldGcvGipGrcGrn	GthHstlelKfdKh	llgnMajMazMi
:	Mostefacil. Touflk			3				
	Mostefaoui, Toufik	Alger				MtiMttNelPorPol	りゃつりゃゃんみゃんらんけい	さらかいらんからんからん
:) } L	Neel, Regis	Venissieux	France	2		MtiMttNelPorPpll		iSauSebSehSu
1 5 1	Neel, Regis Observ. De Brno	Venissieux Brno	France Switzerland	2		MtiMttNelPorPpll SzlTrlTslVczVgl		iSauSebSchSu'
:) } L	Neel, Regis	Venissieux	France	2		· ·	Via,	iSauScbSchSu esents both :

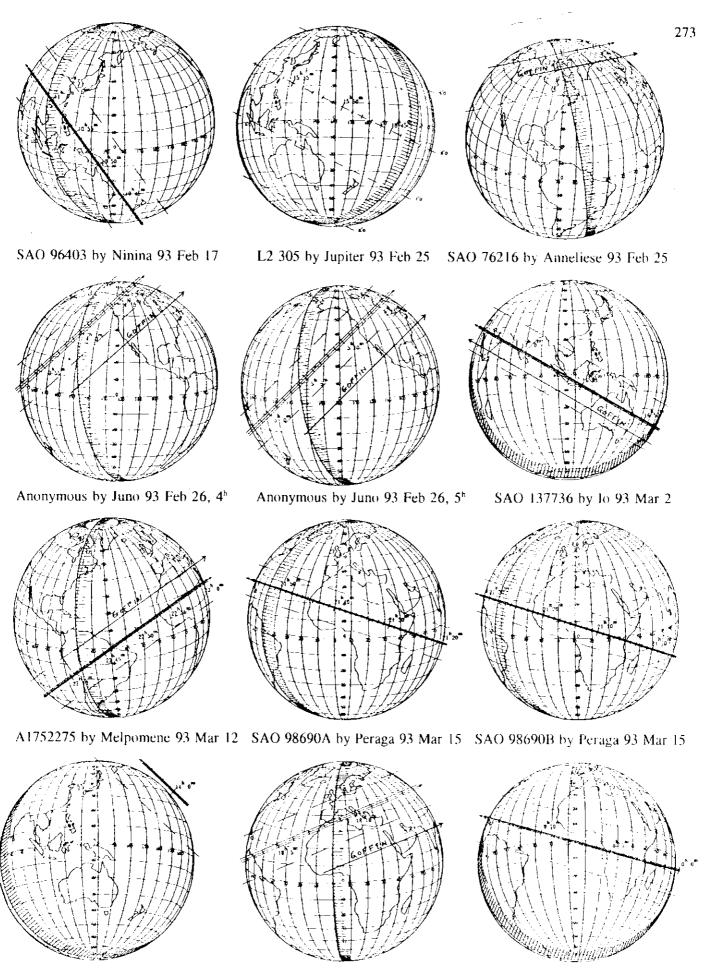
EARLY NEWS ABOUT LAST DECEMBER'S TOTAL LUNAR ECLIPSE

NOTES (Cont.) (Jul-Dec 1990):

- Sep 24 (19) Fortuna. See [O.N.5(4),p.93].
 This is a corrected date.
- 5. Sep 30 Patientia. After playing back his tape on Nov 07, Dik realized that he may have seen somthing. His eyepiece had fogged, but he recorded a "definite" fading of the star "by half its brightness." The times obtained: D at 15:25:26.8 and R at 15:28:04.7.
- 6. Oct 22 Eunike. Blk recorded a D at 18:03:07.6 and a R at 18:03:23.4 (No PEs applied) from Long. E 145° 29' 37.4", Lat. S 37° 22' 36.1", Elev. 260 meters. This shifts the nominal path about 0.7 arcsec south.
- 7. Oct 22 Juewa. 30 observers (BgsBilBnnBul CllDssErnGipHokKbjKlmKrtLgnLntLooMrxPgkPnn ShkSplSprSzaTslTsoVbgVglVsoWbbWhkWls) from 25 stations moonitored this event. Wbb reported a positive event, with no other information, and Dss may have seen something.
- Nov 17 Kleopatra. Harold Reitsema recorded an 11.0 sec. occultation from Dillon, Colorado.
- Dec 09 Interamnia. See [O.N. 5(4), p.93].
 Correction: The observation of AGK# +18°0627 on Mar 13, 1990 [O.N.5(4),p.93] by David Dunham was from Westminster, Maryland.

David W. Dunham

Observers in the northeastern USA and southeastern Canada had a good view of this dark eclipse. Skies were mostly overcast in other areas with large numbers of observers, including the central USA, northwestern Europe, and southern Africa. The eclipse was dark, but there were large variations in brightness within the umbra. Ton Shoenmaker in the Canary Islands and I. observing from Sperry Observatory in Union, NJ, both estimated that the Moon as a whole was about 2nd magnitude near mid-totality, which compares with estimates as low as 4th magnitude for the 1963 December 30th eclipse that was blackened by the recent eruption of Mt. Agung. Although Pinatubo was a larger eruption, apparently the 1.5 years since its eruption has allowed much of the upper-level material that darkens eclipses to settle out of the atmosphere. If the eclipse had been as dark as the one at the end of 1963, it might have been possible to time large numbers of occultations of 13th-magnitude stars. As it was, most 11th-magnitude stars just merged with the Moon's edge without sharp events, like a bright-limb occultation. That was my experience using my 20-cm telescope with image-intensified video in New Jersey, and also Wayne Warren's experience observing visually with a 40-cm Cassegrain telescope at the Goddard Space Flight Center Optical Facility. each timed about 15 occultations of mostly 9th and 10th-magnitude stars. The geometry of the star field for us was such that most of the events were reappearances. I had hoped to use the 61-cm telescope at Sperry Observatory, but other workers there and I were unable to make a practical mechanical connection of my video camera/intensifier with the telescope, so I used my own telescope. The only graze that I now know was observed during the eclipse was the northern-limit one of 7.8-mag. SAO 77019, recorded by Harold Povenmire and a few other observers near Melbourne, FL. Relatively large telescopes were used due to low altitude and some thin cirrus clouds.



TOTA

The International Occultation Timing Association was established to encourage and facilitate the observation of occultations and eclipses. It provides predictions for grazing occultations of stars by the Moon and predictions for occultations of stars by asteroids and planets, information on observing equipment and techniques, and reports to the members of observations made. IOTA is a tax-exempt organization under section 509(a)(2) of the (USA) Internal Revenue Code, and is incorporated in the state of

The ON is the IOTA newsletter and is published approximately four times a year. It is also available separately to non-members.

The officers of IOTA are:

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Executive Vice President	Paul Maley
Executive Secretary	Rocky Harper
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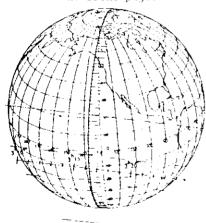
Addresses, membership and subscription rates, and information on where to write for predictions are found on the front page.

The Dunhams maintain the occultation information line at 301-474-4945. Messages may also be left at that number. When updates become available for asteroidal occultations in the central U.S.A., che information can also be obtained from either 708-259-2376 (Chicago) or 713-488-6871 (Houston).

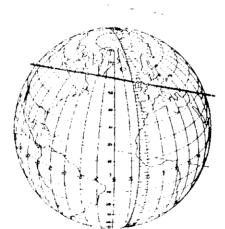
Observers from Europe and the British isles should join IOTA/ES, sending DM 40.-- to the account IOTA/ES; Bartold-Knaust Strasse 8; 3000 Hannover 91; Postgiro Hannover 555 829 - 303; bank-code-number (Bankleitzahl) 250 100 30. Full membership in IOTA/ES includes the supplement for European observers (total and grazing occultations) and minor planet occultation data, including last-minute predictions, when available.

The addresses for IOTA/ES are:

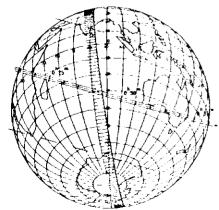
Eberhard Bredner	Hans-Joachim Bode						
Ginsterweg 14	Bartold-Knaust-Str. 8						
D-W-4730 Ahlen 4	(Dolberg) D-W-3000 Hannover 91						
Germany	Germany						
Telephones (49- for	Germany, or 0- within Germany)						
2388-3658	-511-424696						
Fax 2381 36770	-511-233112						



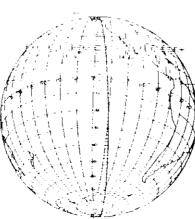
SAO 79187 by Mars 93 Mar 27

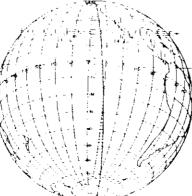


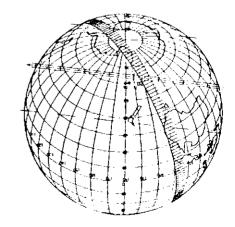
Anonymous by P/Sm-Wm-1 Mar 27 SAO 155934 by Interamnia Apr 1



SAO 211153 by Hektor 93 Apr 2







C23° 14467 by Themis Apr 2 SAO 230574 by Euphrosyne 93 Apr 6